

Structure Growth In Dynamical Dark Energy with Neural ODEs

Peter Taylor



What is driving cosmic
acceleration?

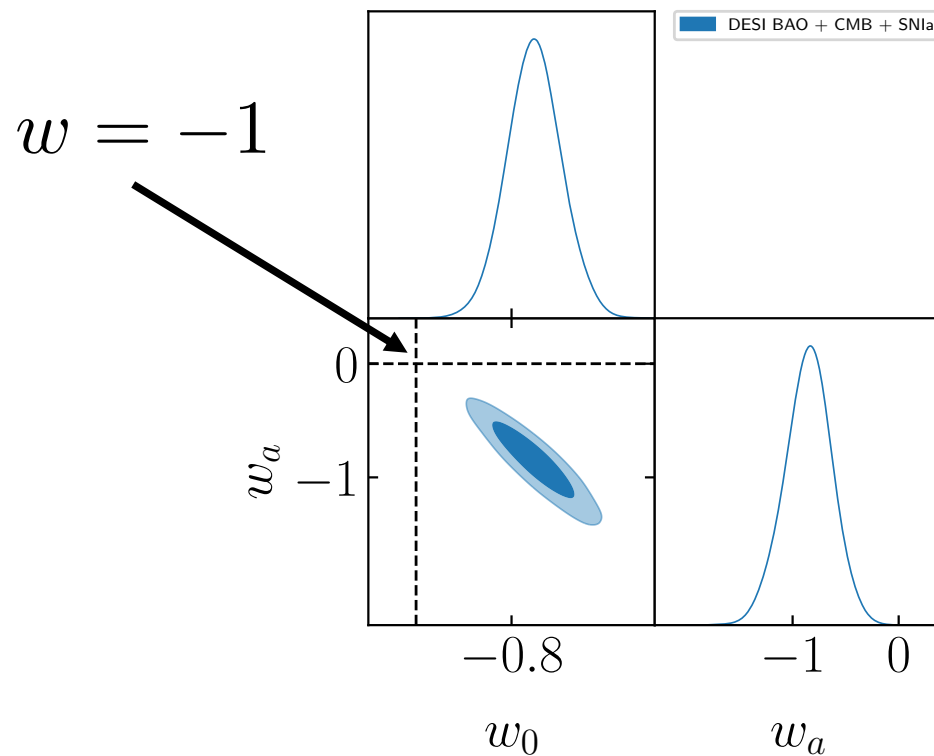
Dark Energy Equation of State

$$w = P/\rho$$

Cosmological Constant (Vacuum Energy)

$$w = -1$$

A Paradigm Shift?

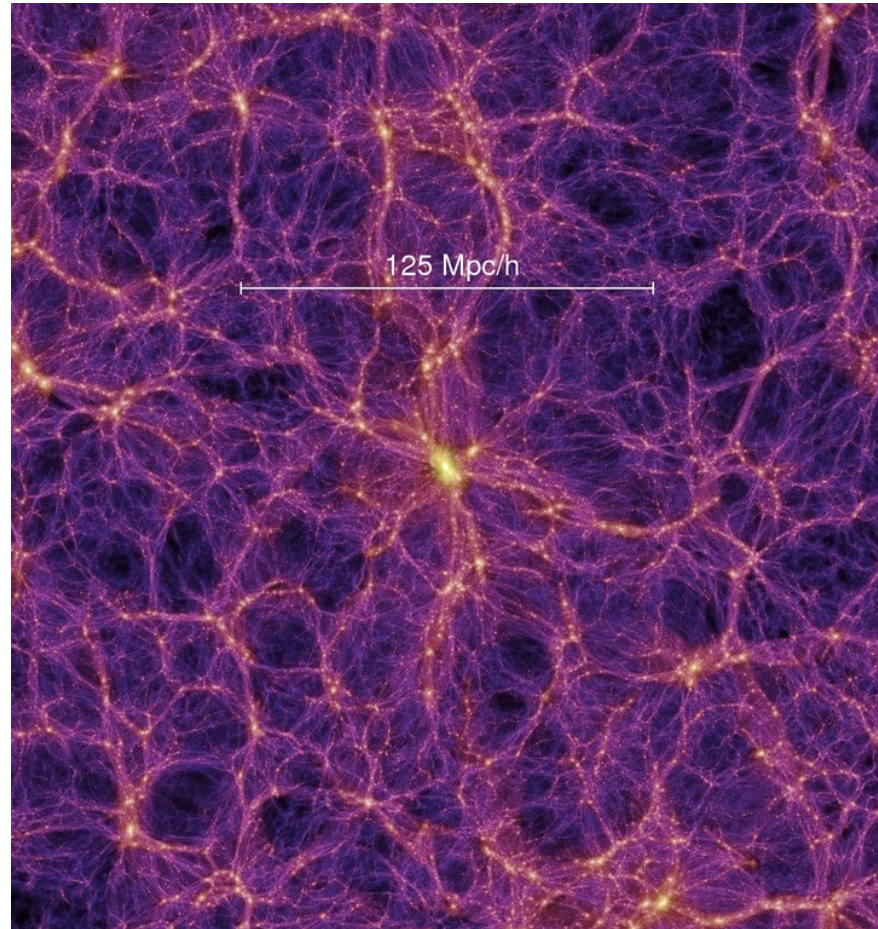


Adapted from DESI Collab. (inc. P Taylor) 25

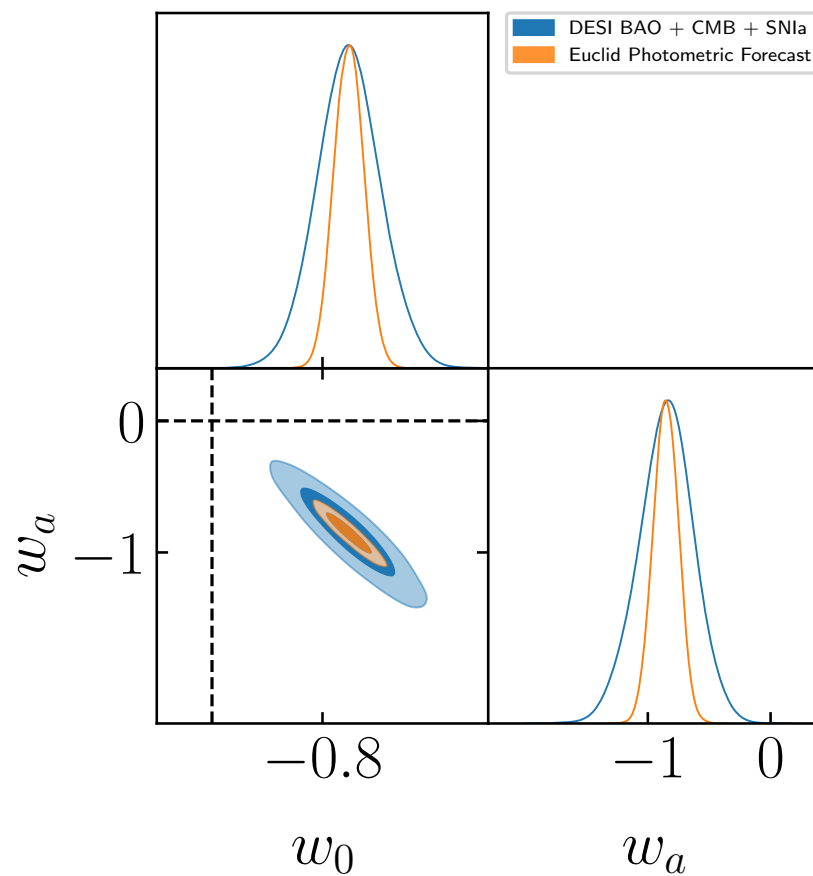


The Next Frontier: Structure Growth

Structure Growth



Improved Precision



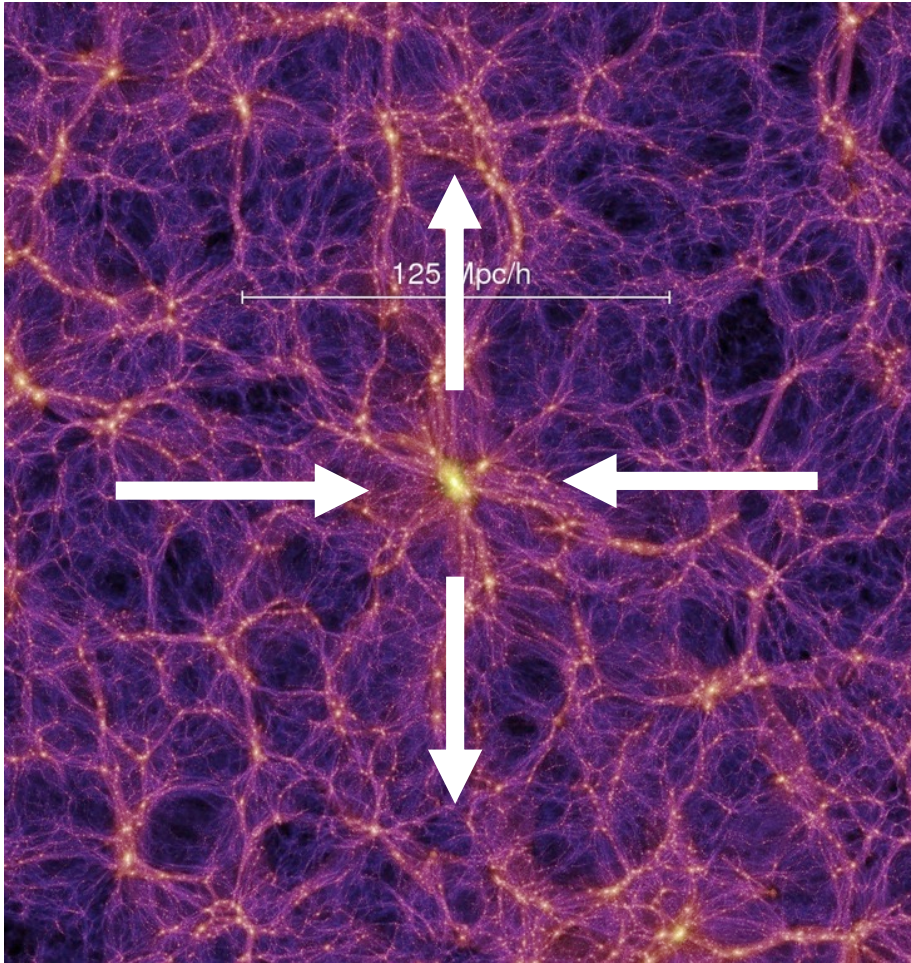


The Problem: No model for
structure growth for arbitrary $w(z)$



Solution: Neural Ordinary Differential Equations (ODE)

The Growth of Structure



grav. collapse

vs.

background expansion

Neural ODE

summary statistic

mean matter density

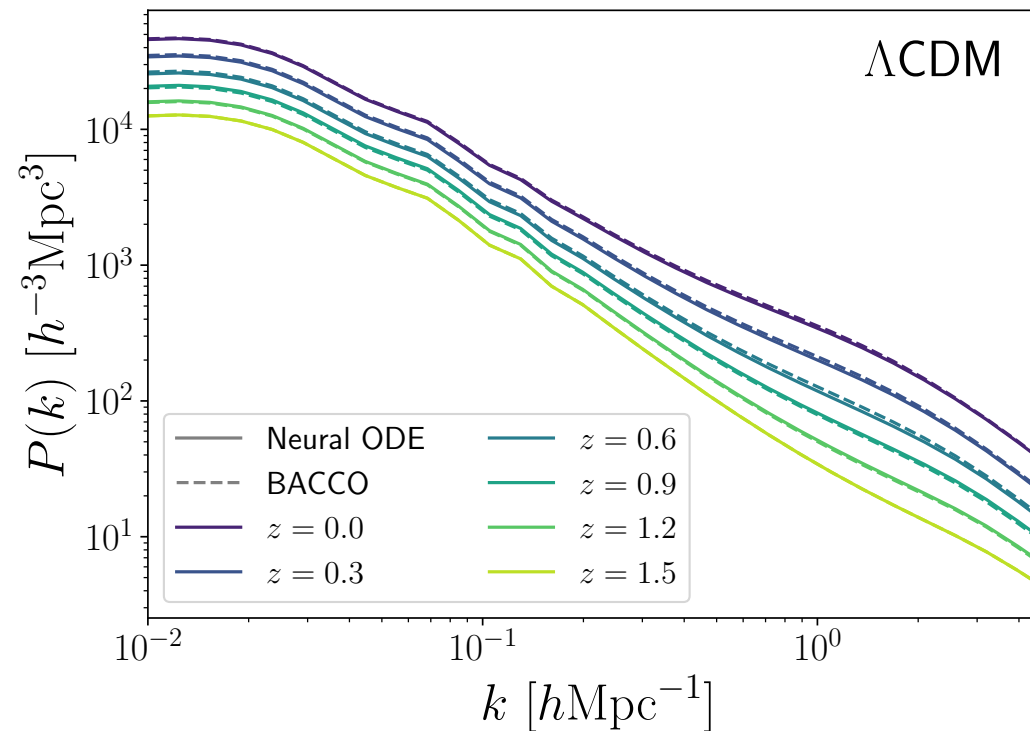
$$\frac{d\mathbf{s}(z)}{dz} = \mathbf{f}(\mathbf{s}(z), H(z), \bar{\rho}_m(z), z)$$

background expansion

A Solution for Structure Growth

$$\mathbf{s}(z) = \mathbf{s}(z_0) + \int_{z_0}^z dz' \mathbf{f}(z')$$

The Power Spectrum (Cosmological Constant)



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Neural ODE

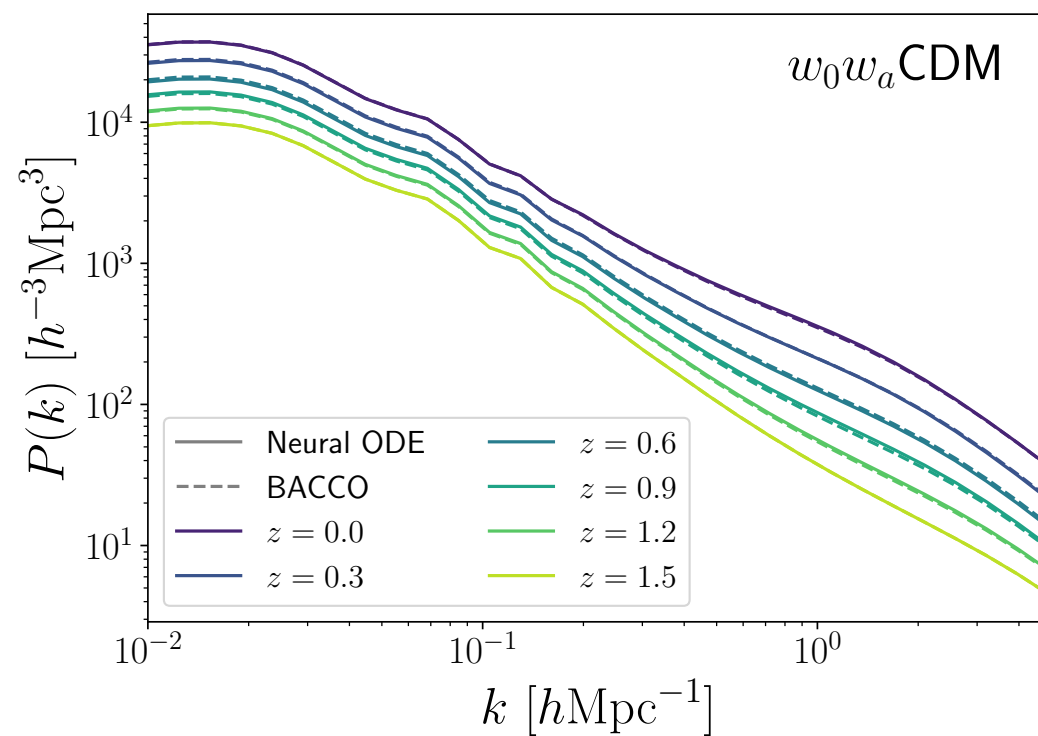
summary statistic

mean matter density

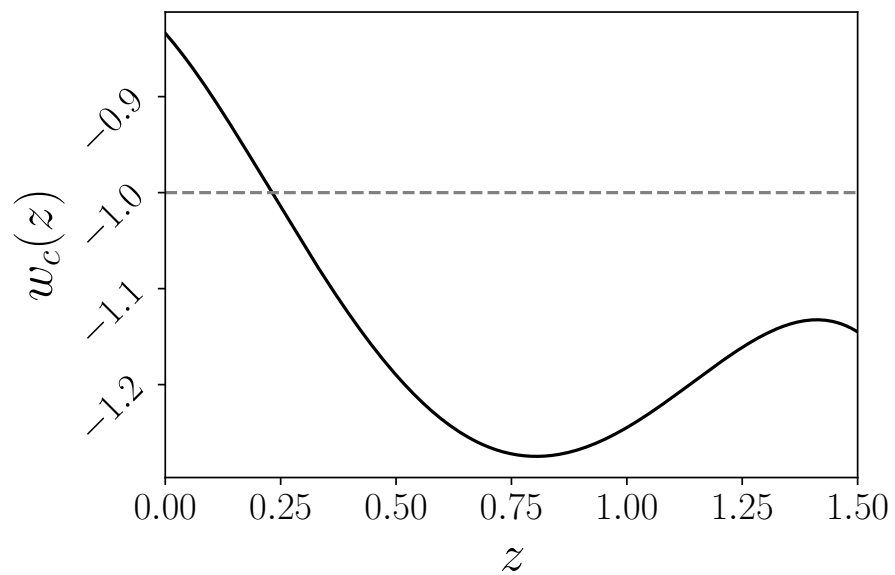
$$\frac{d\mathbf{s}(z)}{dz} = \mathbf{f}(\mathbf{s}(z), H(z), \bar{\rho}_m(z), z)$$

background expansion

Model Generalizes Across DE Models

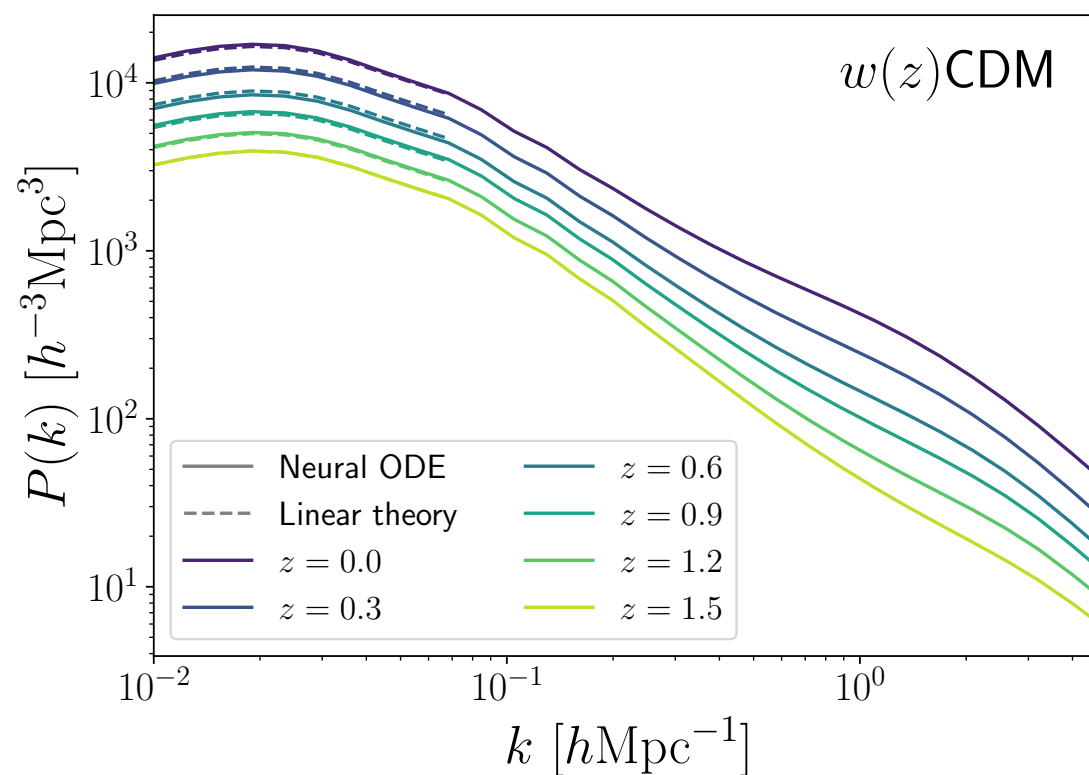


Custom $w(z)$



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Structure Growth in For Custom $w(z)$



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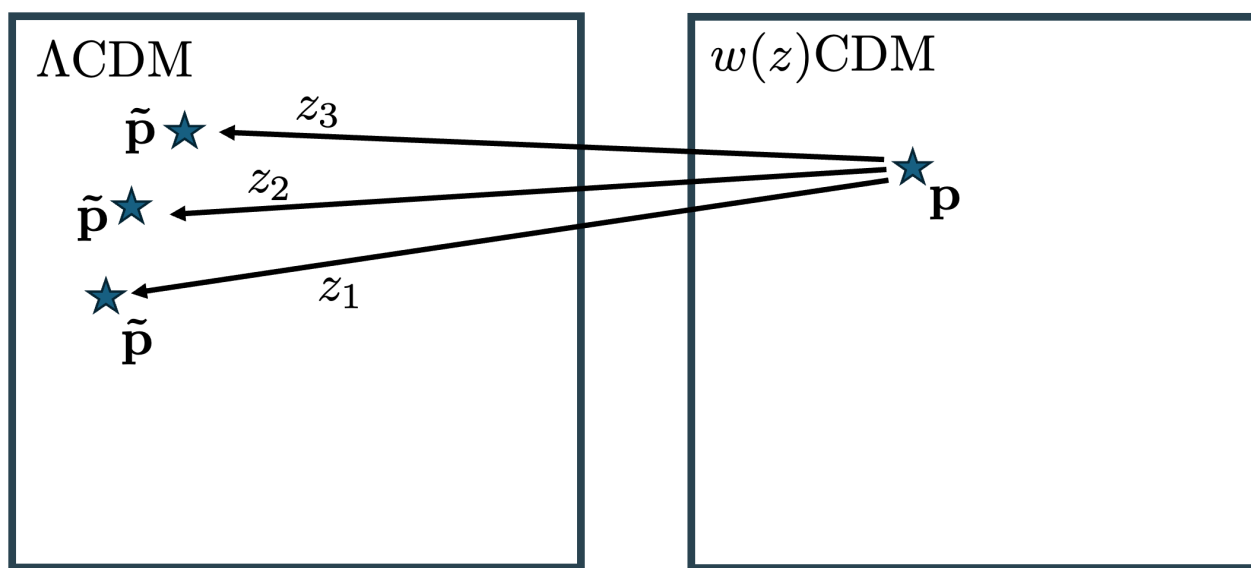
Conclusion

- **Structure growth provide additional constraining power on DE.**
- **Neural ODEs unlock structure growth across dynamical dark energy model space.**
- **Future Plans: Additional statistics + application to data.**



Extras

Model Generalizes Across DE Models



$$\bar{\rho}_m(\mathbf{p}), H(\mathbf{p}), P(k; \mathbf{p}) = \bar{\rho}_m(\tilde{\mathbf{p}}), H(\tilde{\mathbf{p}}), P(k; \tilde{\mathbf{p}})$$