

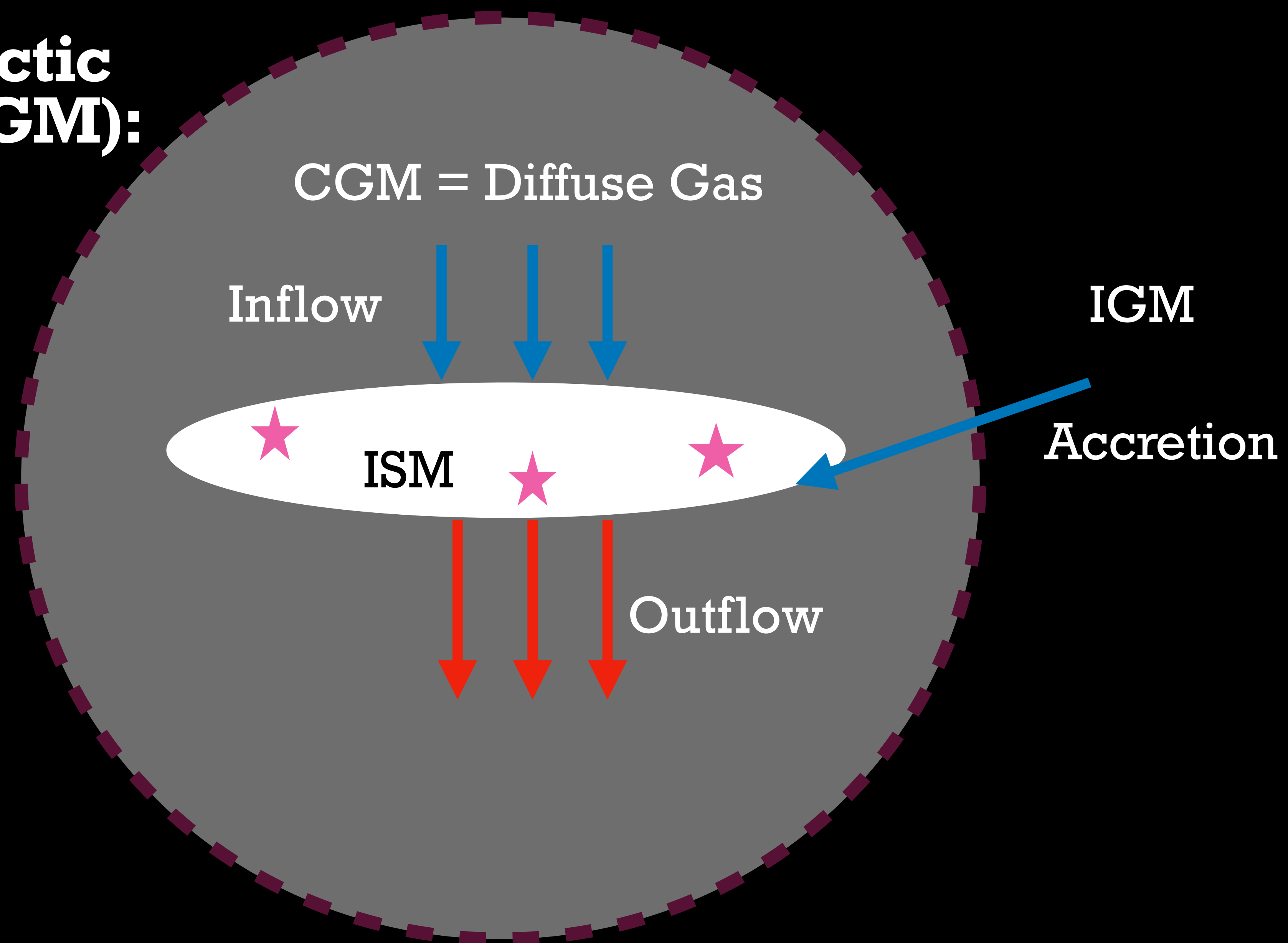
Cosmic Rays: The Hidden Architects of the Circumgalactic Medium

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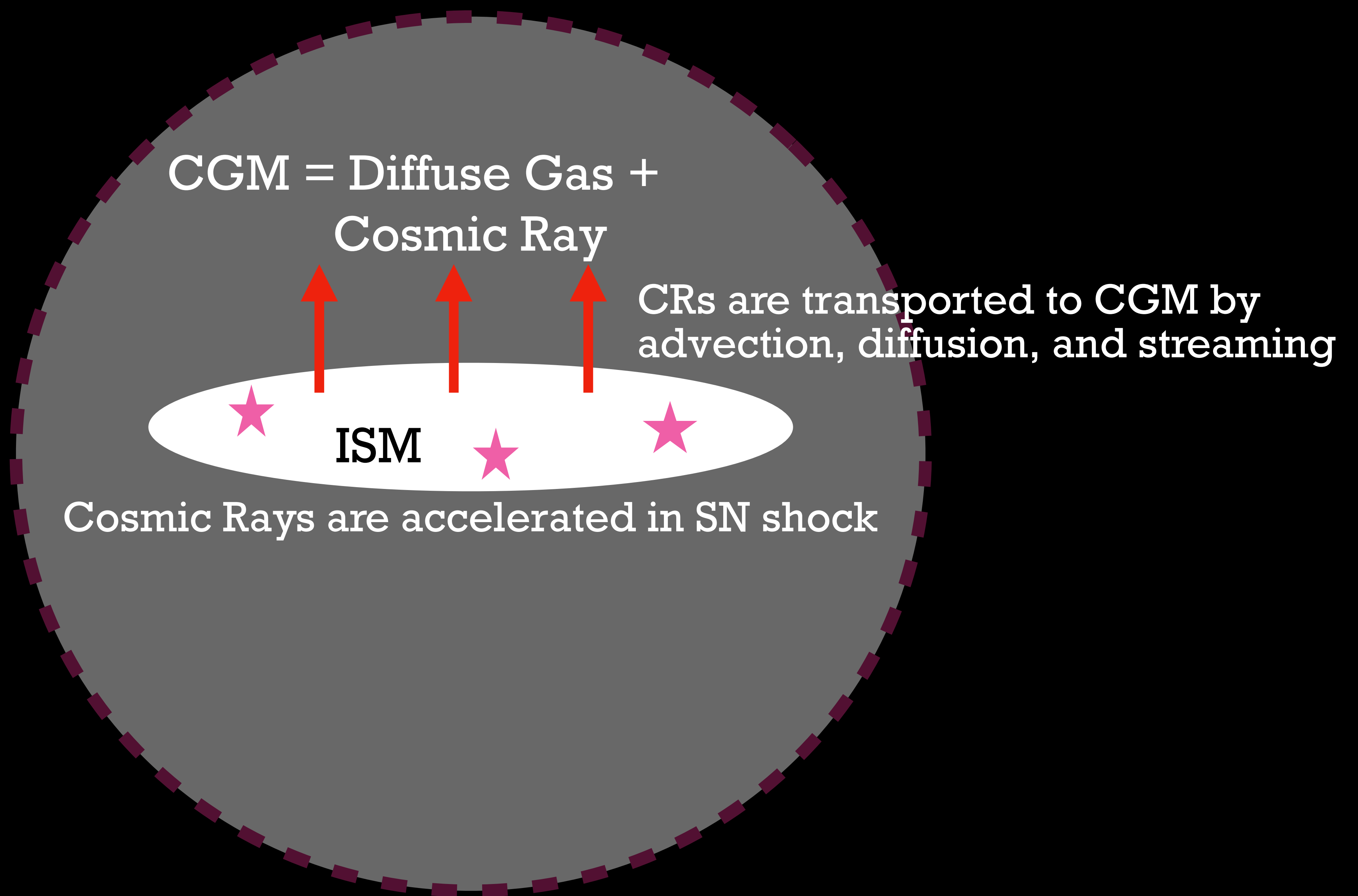


With Kung-Yi Su, Stephanie Tonnesen

Circumgalactic Medium (CGM):



CGM is important for galaxy evolution





The diagram consists of a large gray circle with a dashed purple border, representing the CGM. Inside this circle is a smaller white horizontal oval representing the ISM. The label 'CGM' is at the top of the gray circle. The label 'ISM' is in the center of the white oval. Three pink stars are positioned horizontally within the white oval: one on the left, one in the center, and one on the right. Below the ISM oval, within the gray circle, is a block of text explaining that CR transport and amount are known in the ISM through various observations.

CGM

ISM

CR transport and the amount of CR
are known
in the ISM by observations
like B/C ratio, gamma-ray emission,
radio synchrotron,
and local measurements.



CGM

ISM

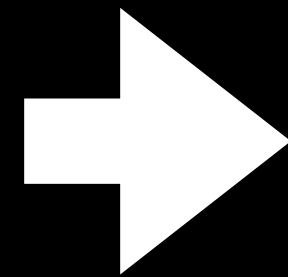
On the contrary, CR transport in CGM is
an open question due to a lack
of observations.

Main Questions:

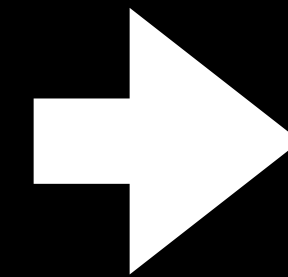
1. How to constrain CR transport in CGM?
2. What are the effects of CR in the CGM?

How to constrain Cosmic rays in the CGM?

CR MODELS in Simulation



Produce different galaxy/
CGM observables

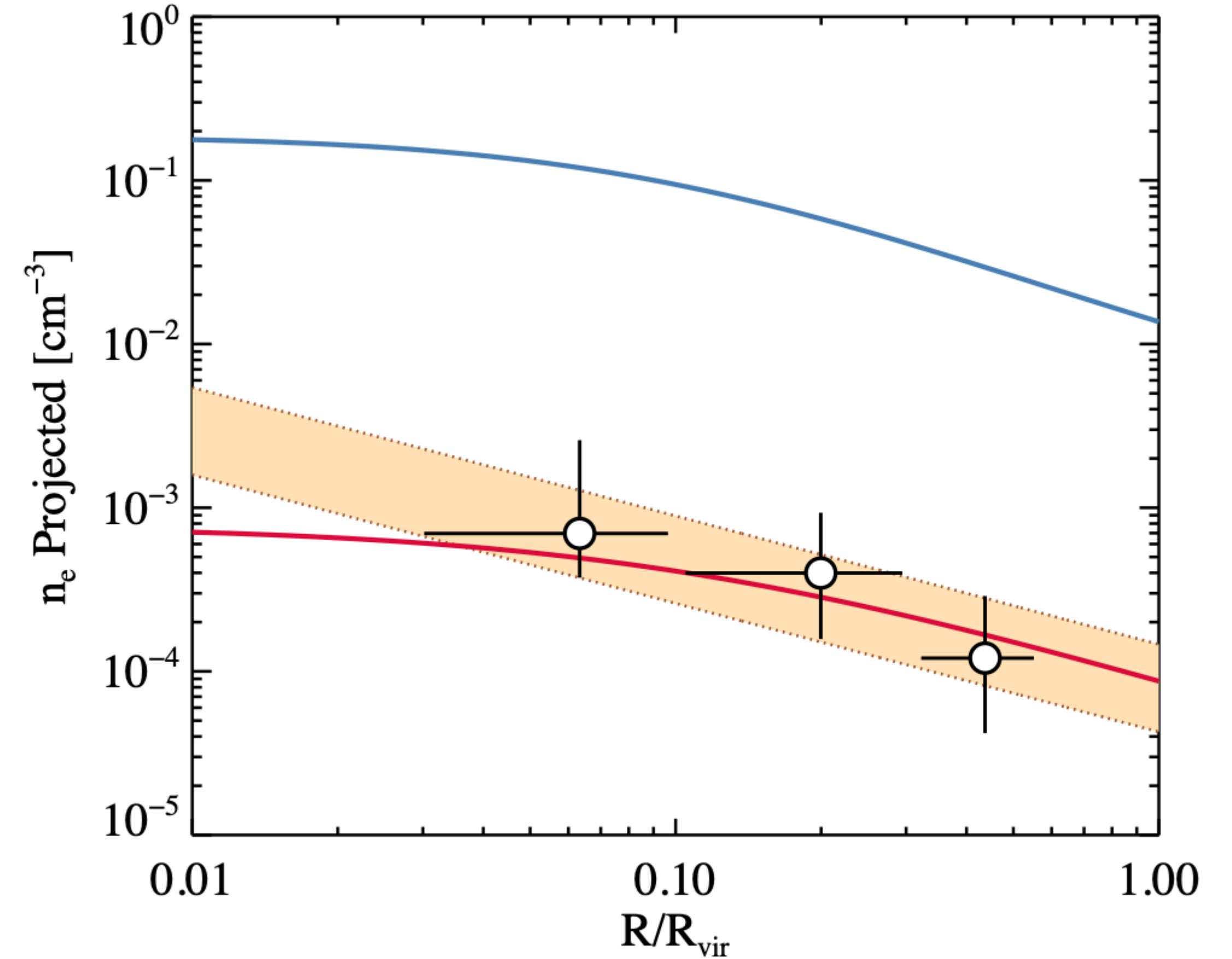
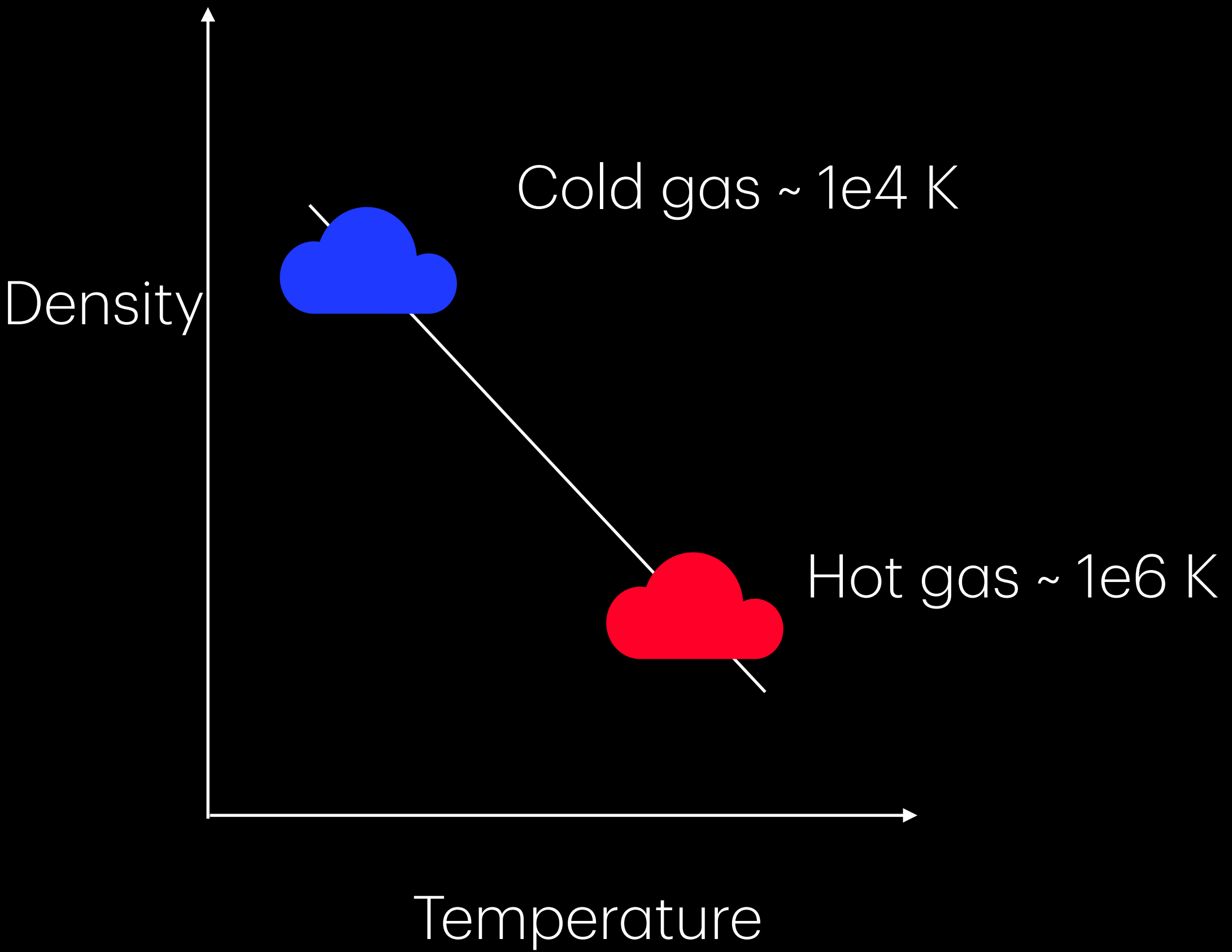


Match with observations

1. Constant Diffusion
 2. Extrinsic Turbulence
 3. Self Confinement
- And so on...

1. Galaxy morphology
2. Different ion column
3. Thermal Pressure
4. Covering Fraction

Cold gas in the CGM



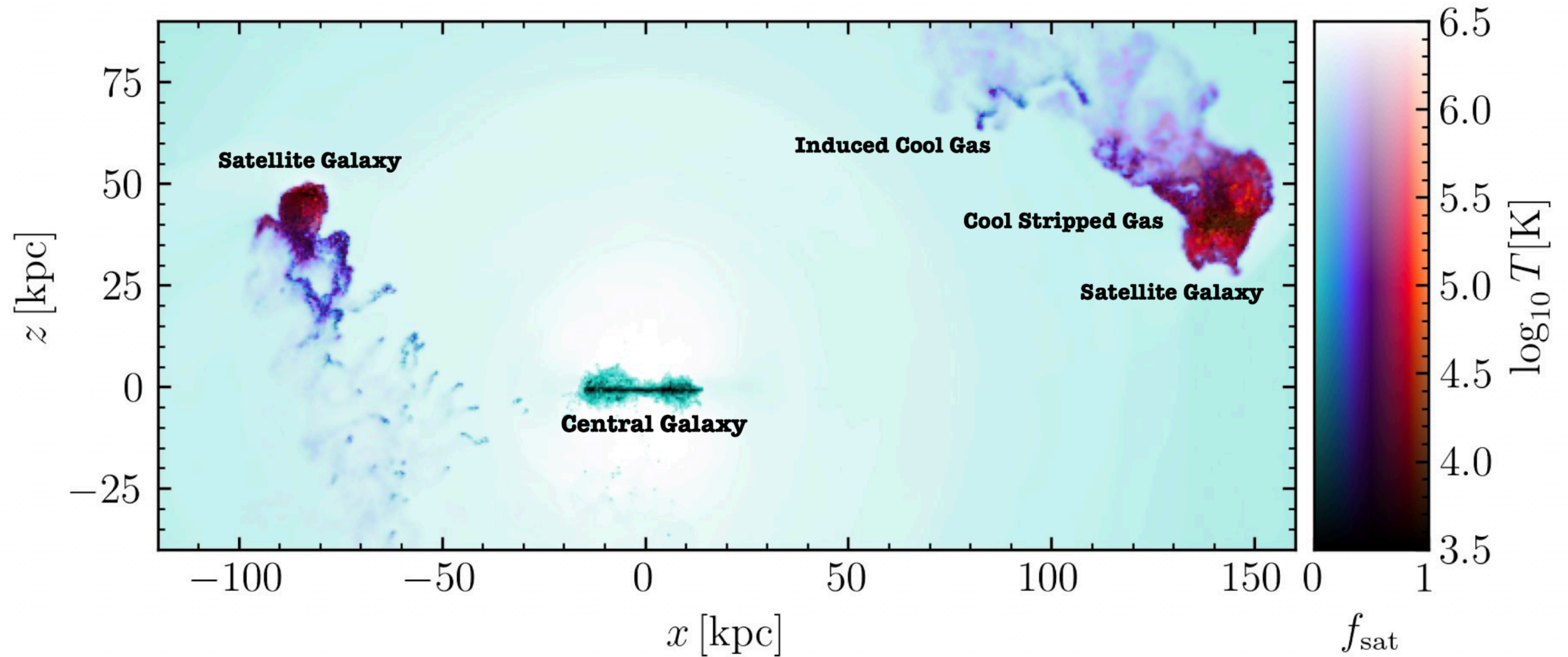
Simulation Setup:

1. Injection of CR by SN in the host and satellite
2. Constant Diffusion + Advection + Streaming



GIZMO, FIRE-2 Feedback

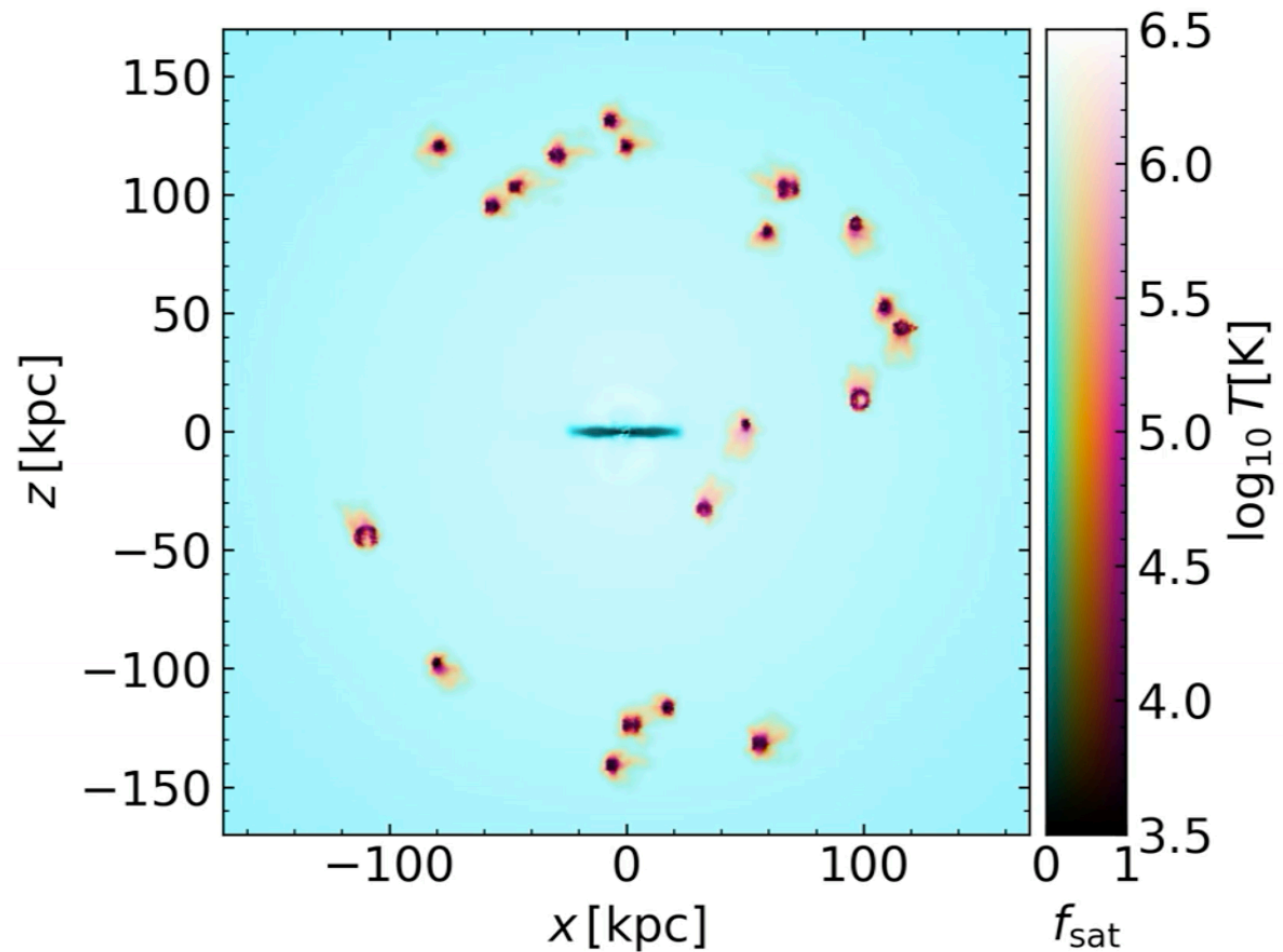
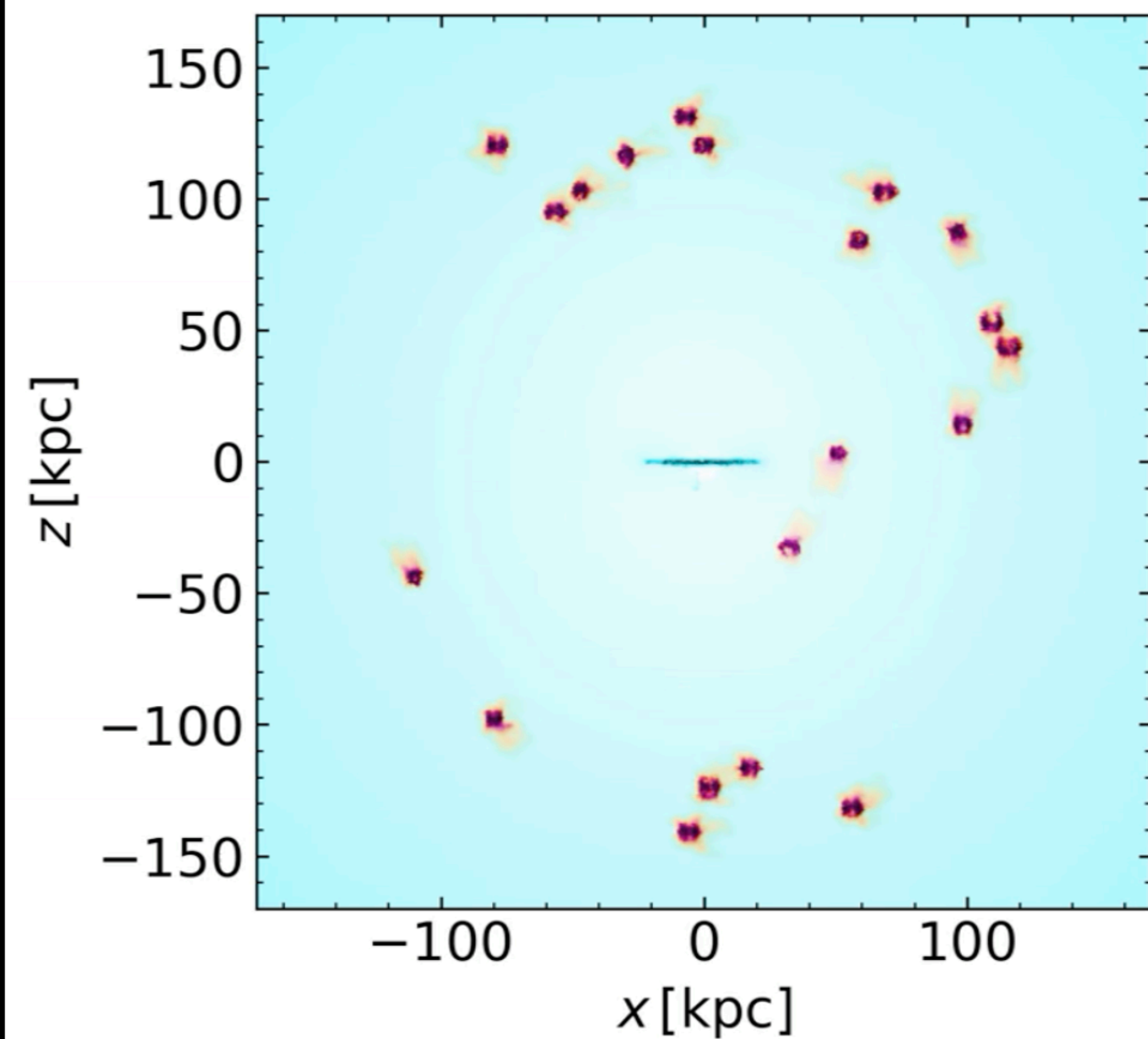
Roy +2025 Submitted in ApJ



Roy +2024

<https://doi.org/10.1093/mnras/stad3142>

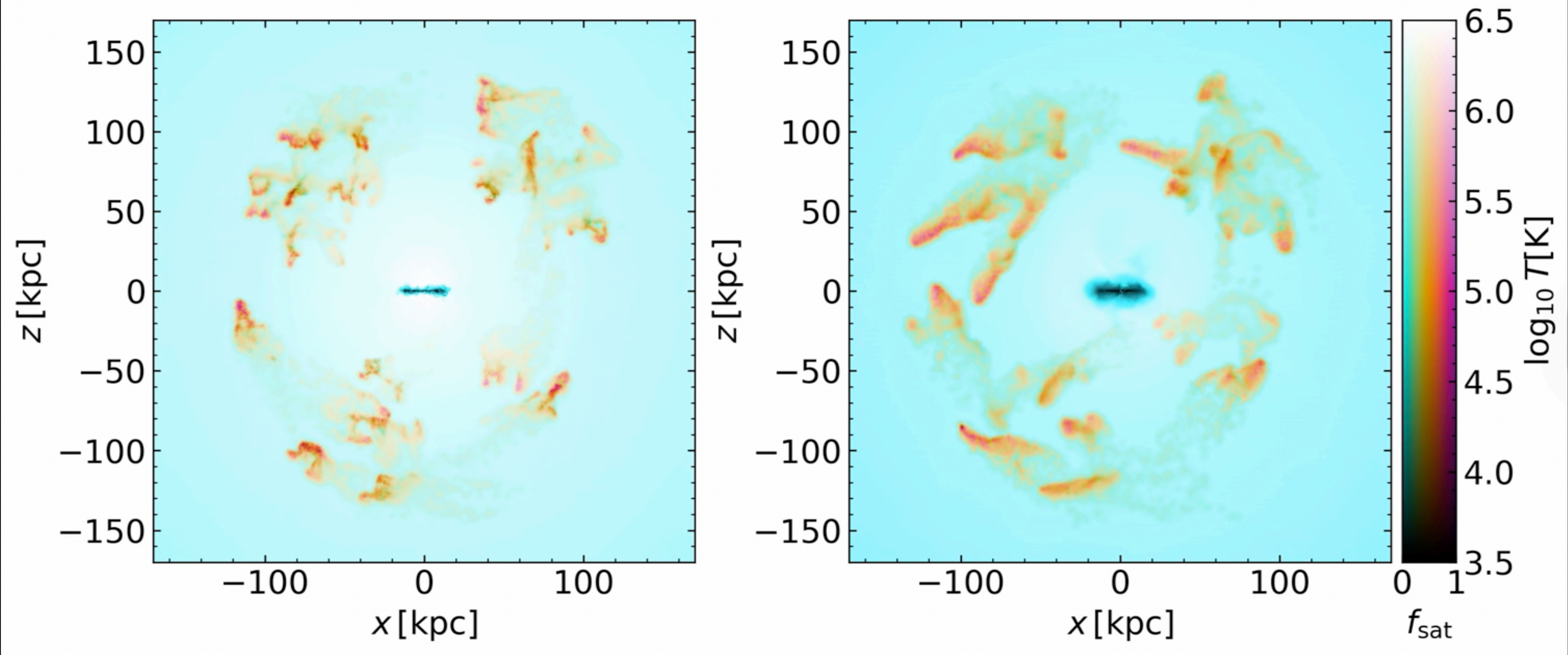
$t = 0$ Gyr



No CR

With CR

Later Time: $t \sim 1 \text{ Gyr}$



No CR

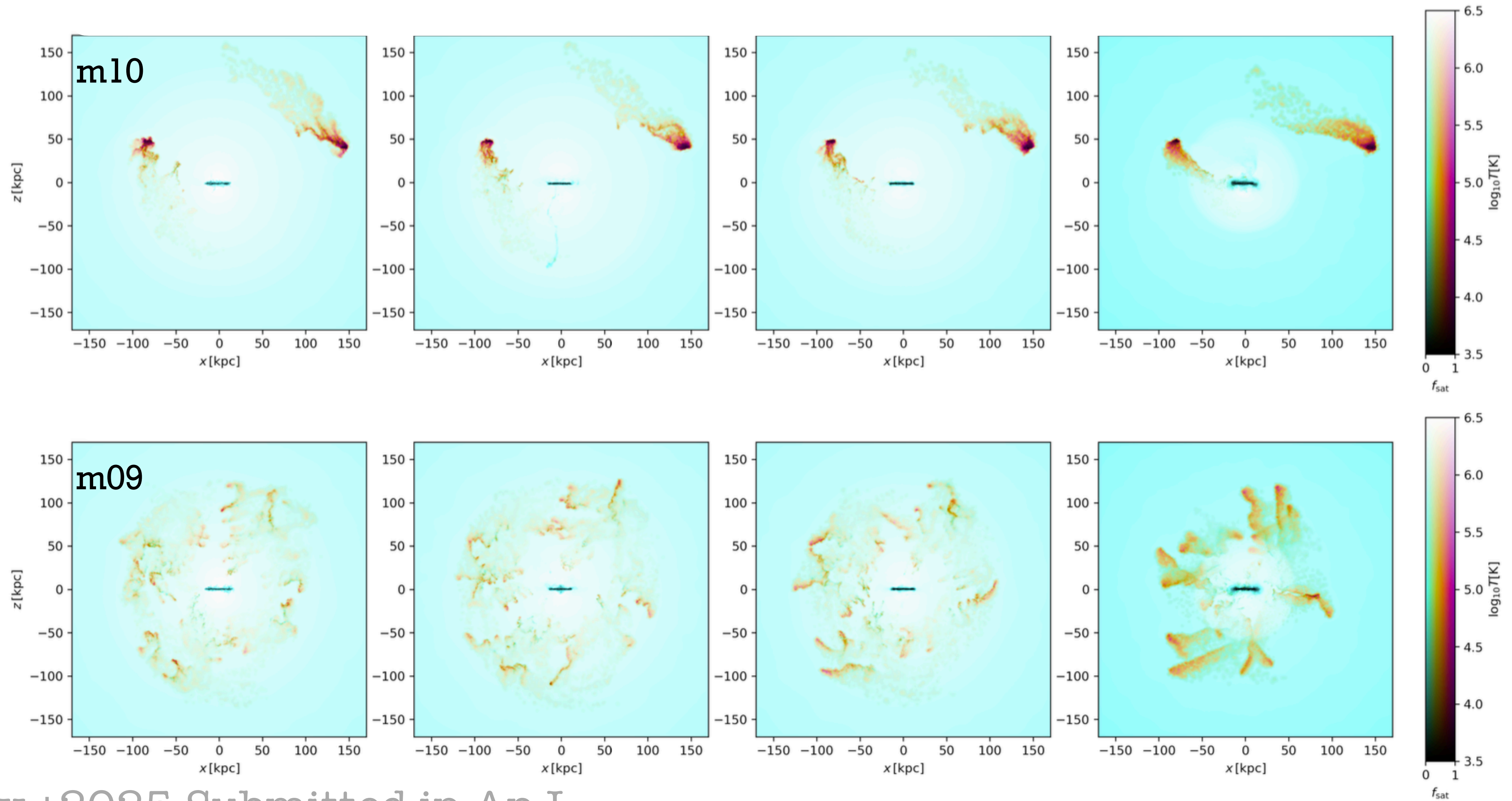
With CR

No CR

Low CR

Mid CR

High CR



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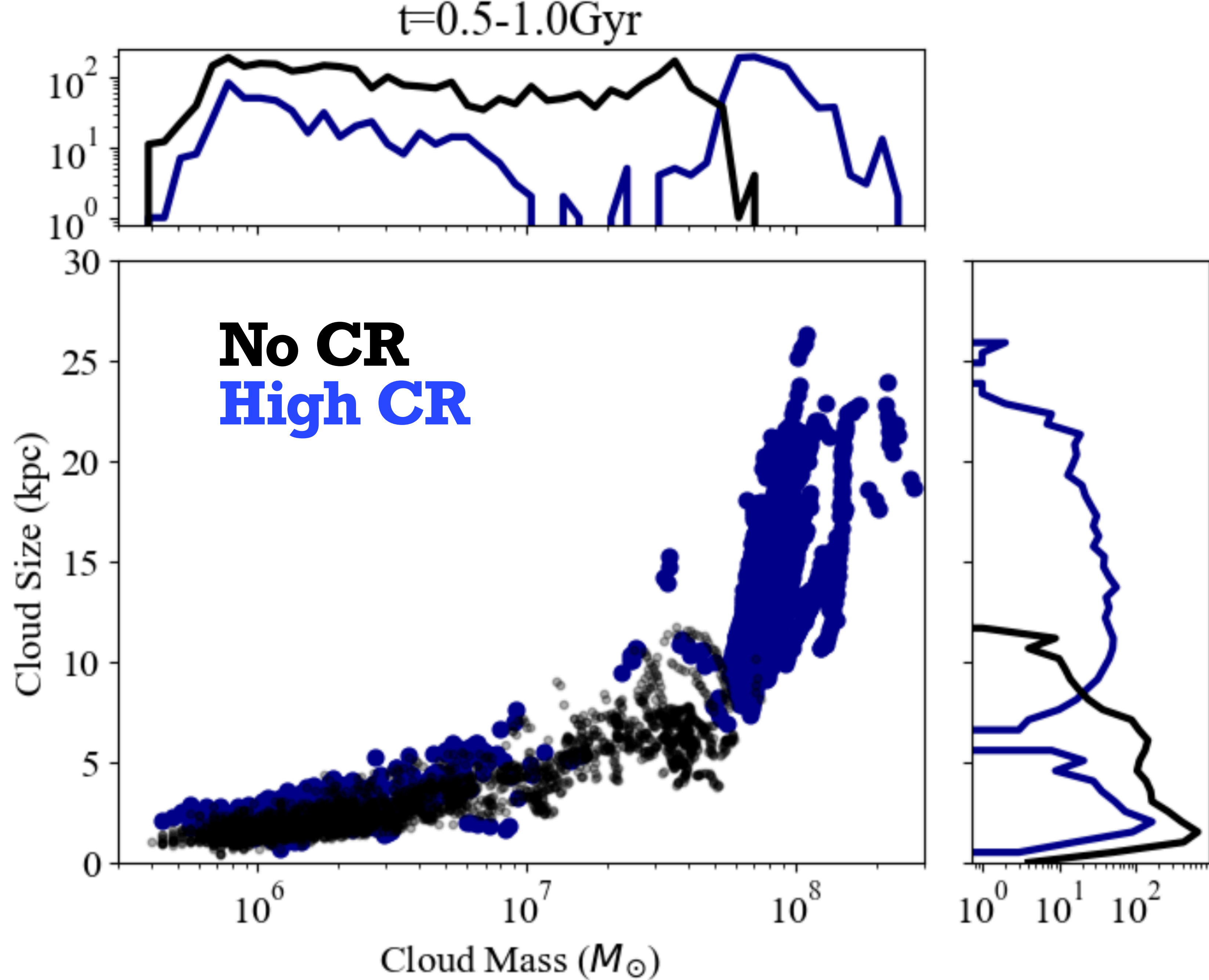


Increasing Cosmic Rays > Increase size of the cloud

Size vs mass:

Larger clouds
with an increase
in CR

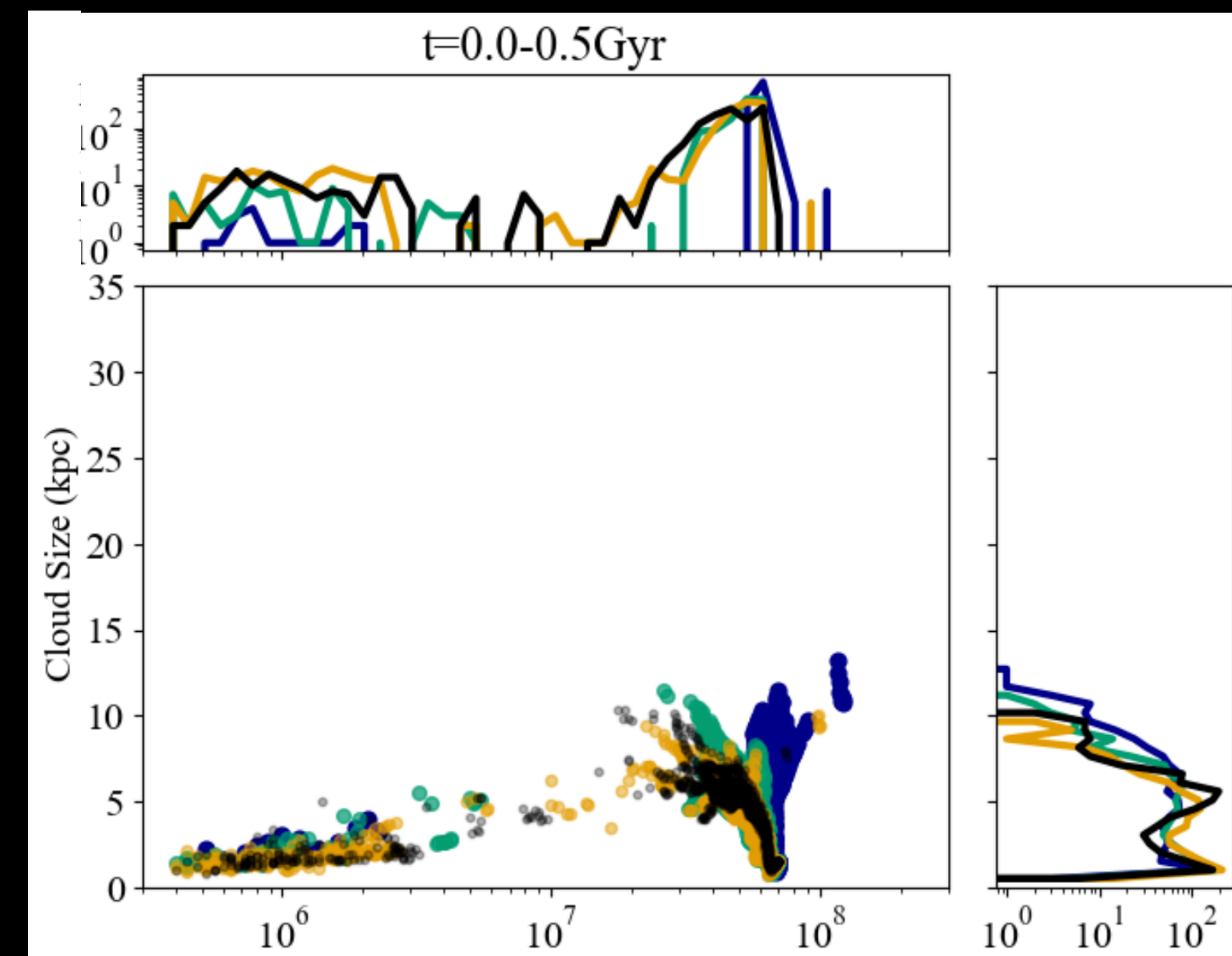
Roy +2025 Submitted in ApJ



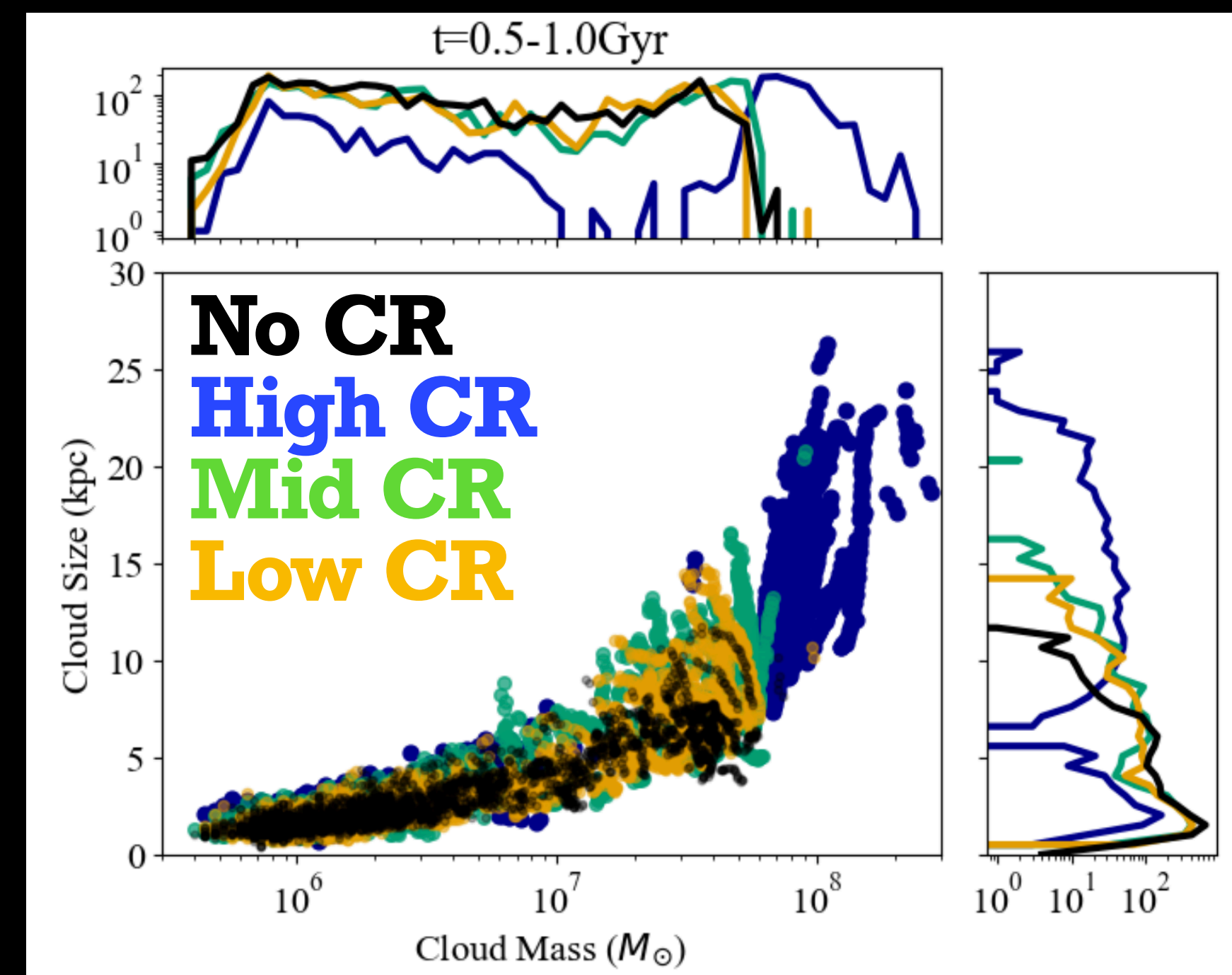
Size vs mass:

Earlier Time

Larger clouds and
cloud mass growth
with an increase
in CR



Later Time



Roy +2025 Submitted in ApJ

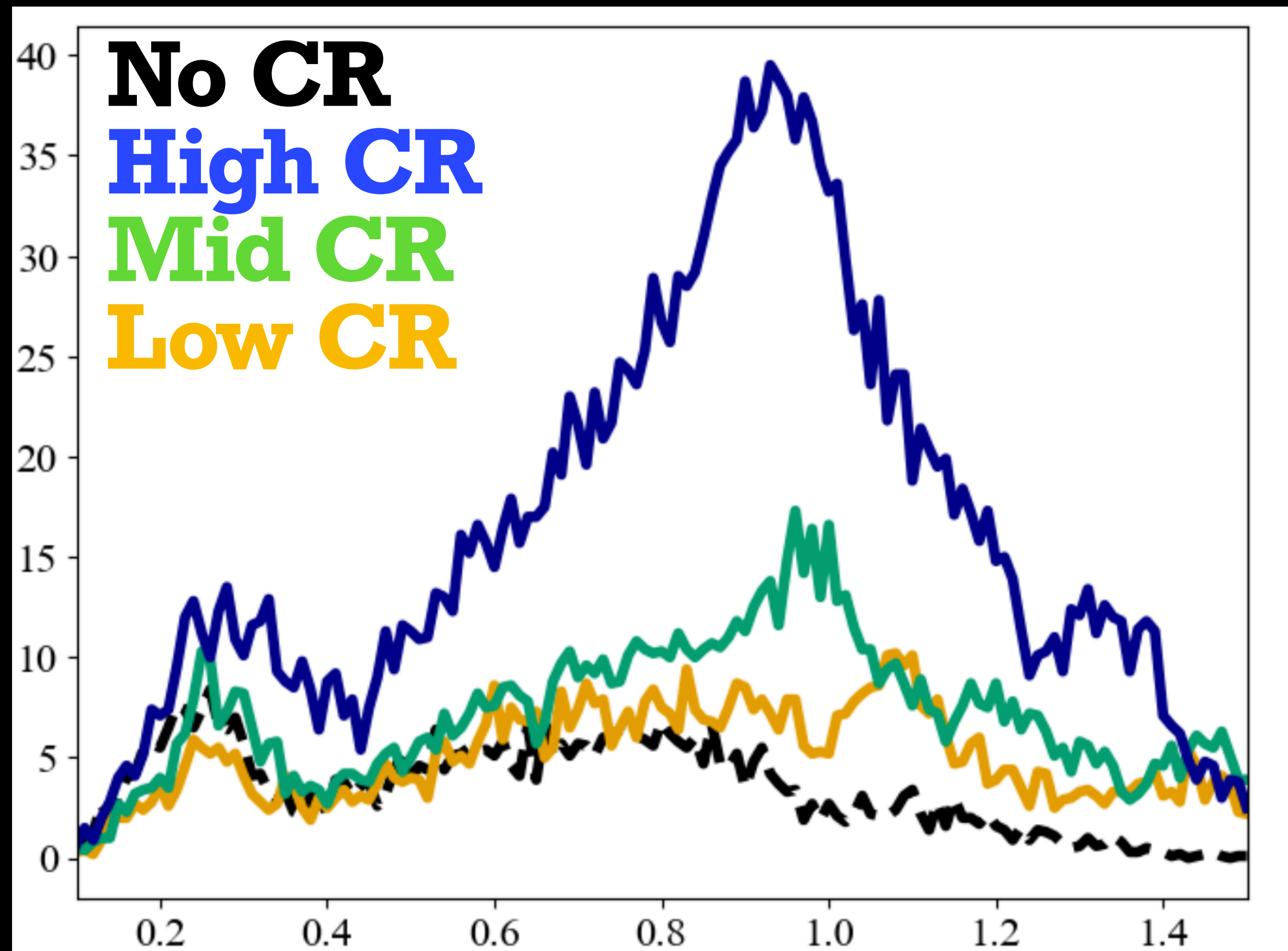
Covering Fraction:

MW

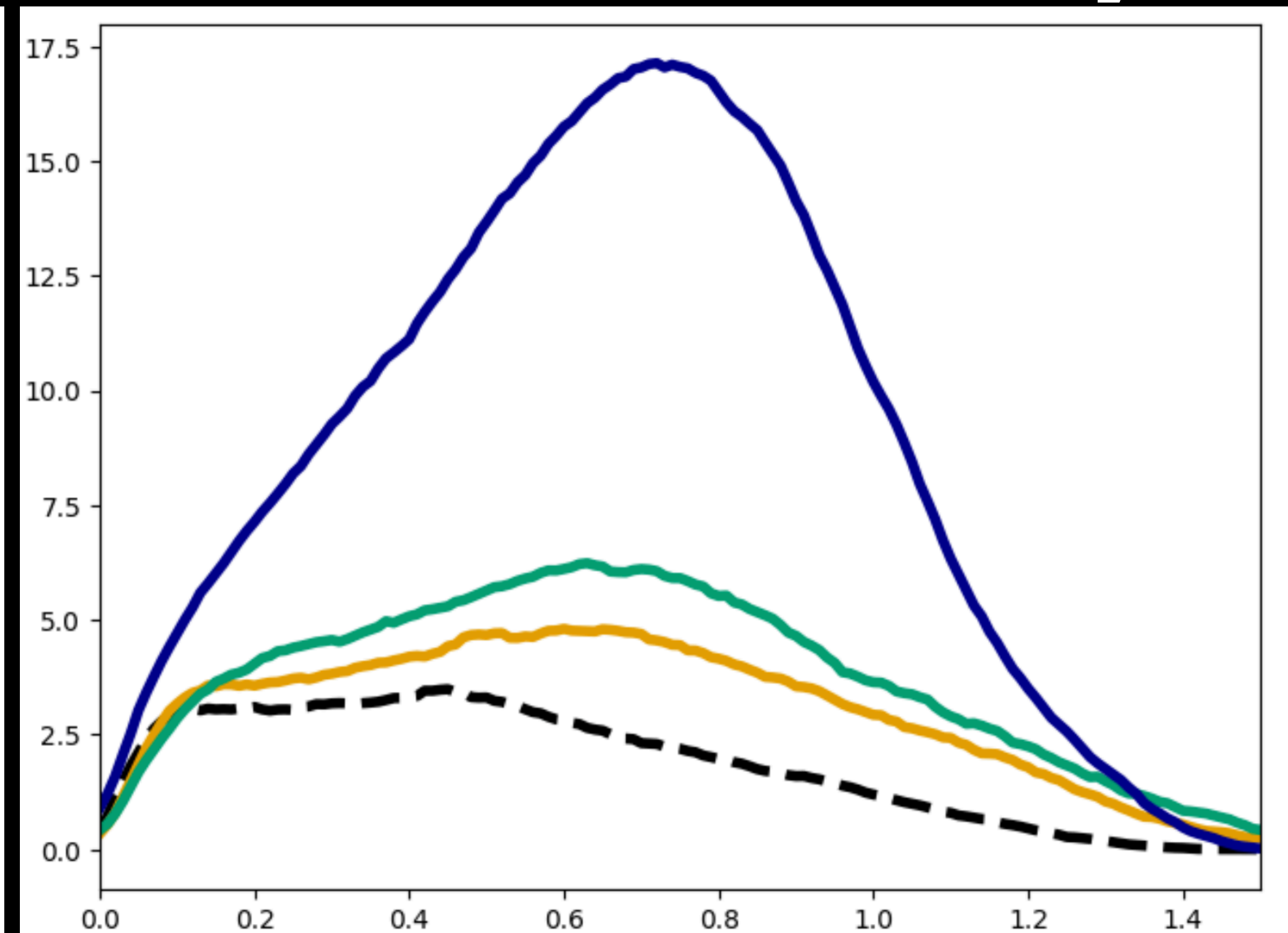
External Galaxy

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Covering Frac (%)



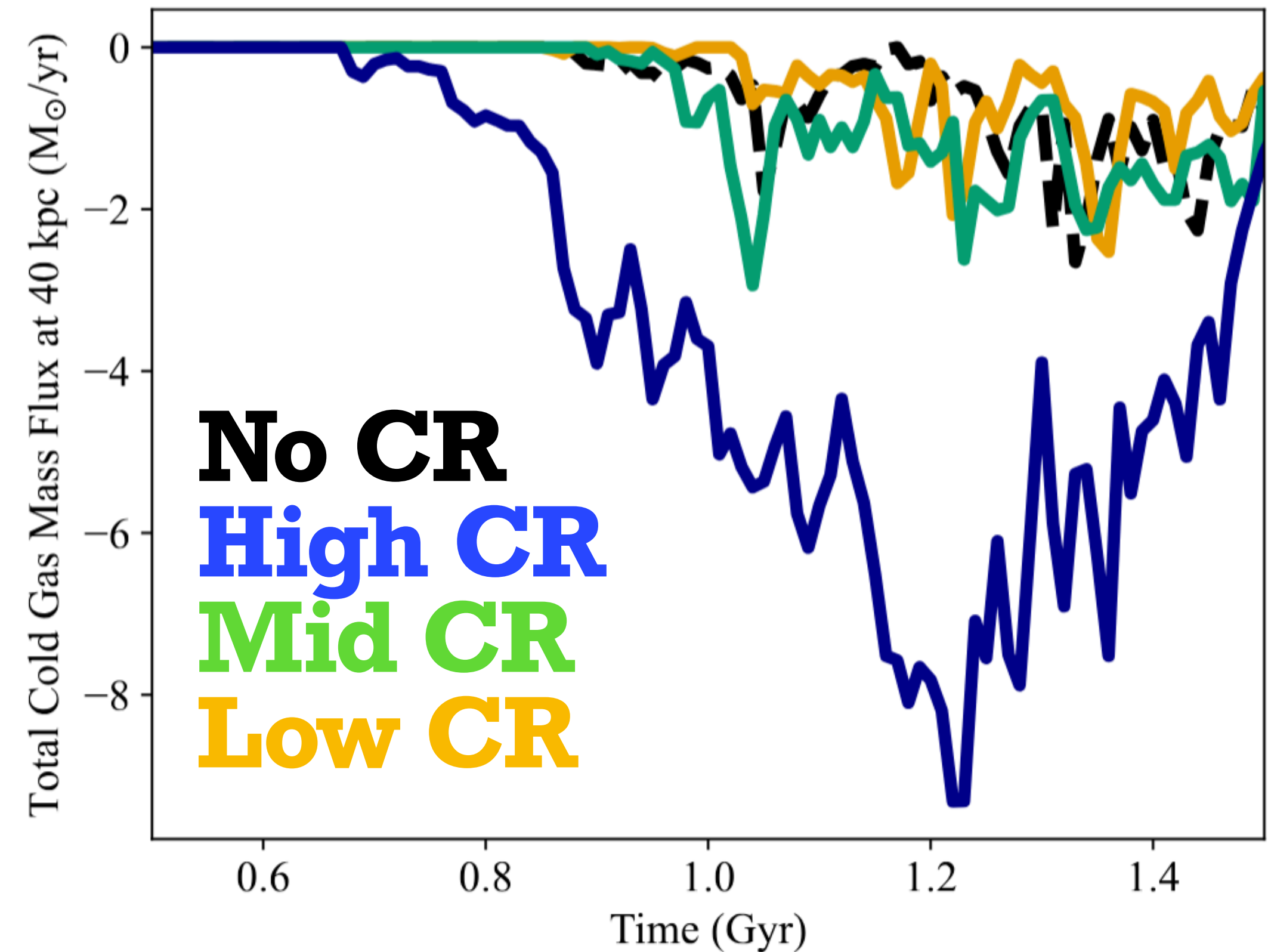
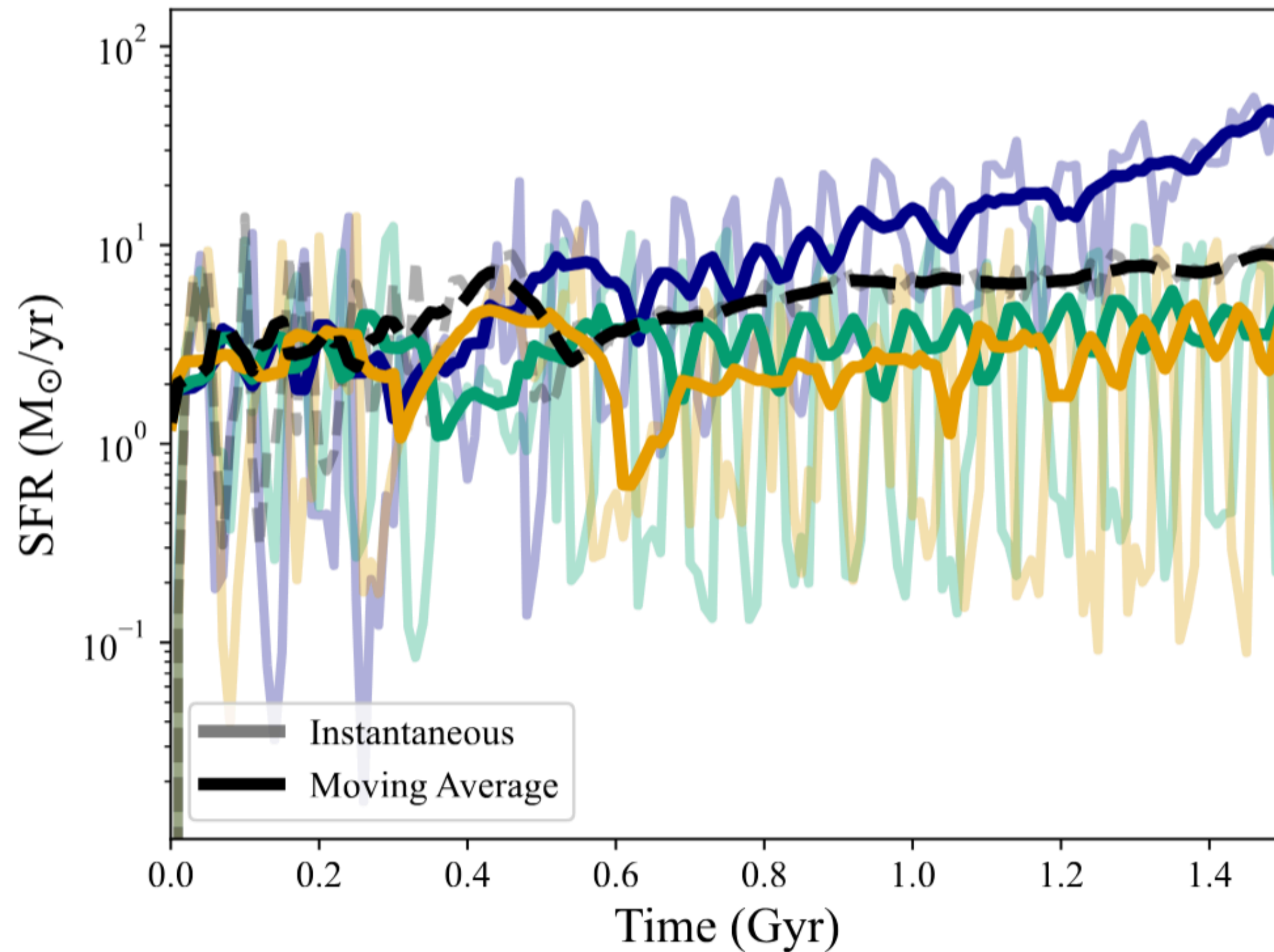
Time (Gyr)



Time (Gyr)

Cosmic rays significantly enhance the covering fraction of cold gas in CGM, particularly in the MW view. It is also true for external galaxies, but in satellite-rich environments like the m09 run

Feeding the galaxy:



Inflow is 3-4 times higher, and SFR is also 2 times higher in the high-CR run than in the no-CR run

Roy +2025 Submitted in ApJ

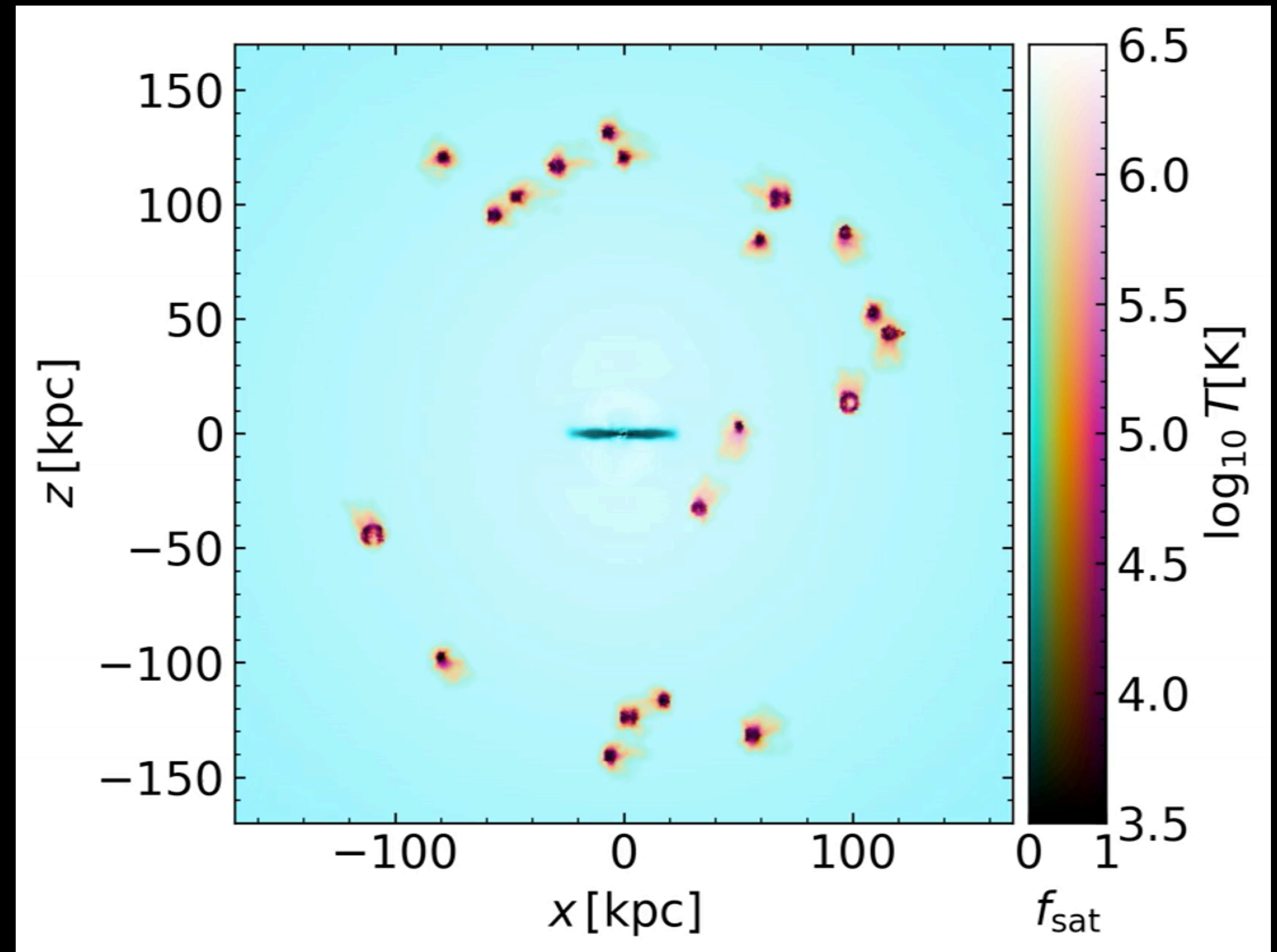
Take Home

Paper I (Roy+2024)

Paper II (Roy+2025 to be Submitted to ApJ)



- ***CR makes the clouds bigger in size which they grow more by mixing layer cooling***
- ***Covering Fraction of the cold gas increases***
- ***Inflow of the cold gas into the central galaxy increases***
- ***SFR of the host galaxy can be boosted at a later time***



Effect of CR transport:

